

### 4.3 Quantized Atom + Quantized Field

Note in semiclassical  $P_{fi} \propto |H_{fi}^I|^2 = |H_{if}^I|^2 \propto P_{if}$

Engineers call the reciprocity. If  $i > f$  this is probability of stimulated emission

$$|H_{if}^I|^2 \propto |\vec{E}_0 \cdot \vec{d}_{if}|^2 = |\vec{E}_0 \cdot \vec{d}_{fi}|^2$$

Hence down  $\downarrow_{if}^i$  and up  $\uparrow_{if}^f$  occur with equal prob

(detailed balance). If  $E_0 \equiv 0$  no transitions

occur. Quantum field  $\Rightarrow |i\rangle \rightarrow |f\rangle \Rightarrow$

possible if  $i > f$  and  $E_0 = 0$ . Vacuum stimulates emission.

In dipole approximation  $e^{i\vec{k} \cdot \vec{r}} \approx 1$  and 2.130  $\Rightarrow$

$$\hat{\vec{E}}(\vec{r}, t) \approx \vec{E}(t) = i E_0 \vec{e} \left[ \underbrace{\hat{a} e^{-i\omega t}}_{\hat{a}(t)} - \underbrace{\hat{a}^\dagger e^{i\omega t}}_{\hat{a}^\dagger(t)} \right]$$

where  $E_0 = \sqrt{\frac{\hbar\omega}{2\epsilon_0 V}}$  and  $\vec{E}_0 \equiv i E_0 \vec{e}$

This is Heisenberg picture where  $\hat{H}^I = \hat{H}_H^I(t)$

Recall (Merzbacher)  $\hat{\Theta}_H(t=0) = \hat{\Theta}_S$  for any

Hermitian operator

$$\Rightarrow \hat{\vec{E}} = \vec{E}_0 [\hat{a} - \hat{a}^\dagger] \quad \text{In Schrodinger}$$

$$\hat{H}_0 = \underbrace{\hat{H}_{\text{atom}}}_{\frac{\hat{p}^2}{2m} - \frac{e^2}{r}} + \underbrace{\hat{H}_{\text{field}}}_{\hbar\omega \hat{a}^\dagger \hat{a}} \Rightarrow \text{atom uncoupled to field}$$

normal ordered - no vacuum

$$\hat{H} = \hat{H}_0 + \hat{H}_I \leftarrow \text{atom-field coupling}$$

$$\hat{H}_I \equiv -\hat{d} \cdot \hat{\mathbf{E}}$$

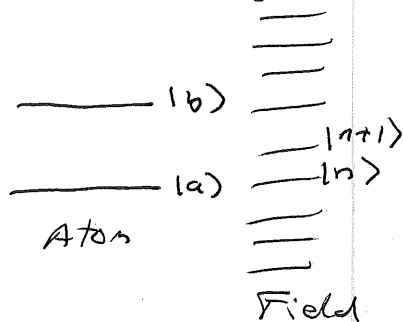
$$= \boxed{-\vec{E}_0 \cdot \hat{d} (\hat{a} - \hat{a}^\dagger)}$$

Hilbert space is tensor product of atom states

$$|a\rangle, |b\rangle \text{ s.t. } \hat{H}_{\text{atom}} |a\rangle = E_a |a\rangle$$

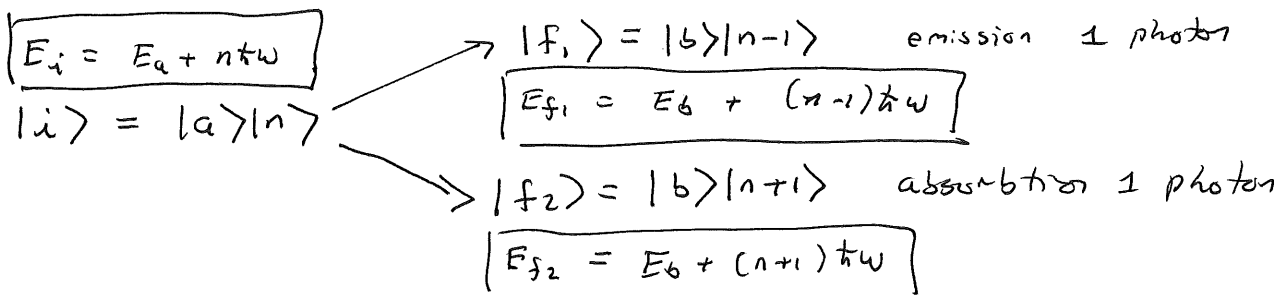
$$\hat{H}_{\text{atom}} |b\rangle = E_b |b\rangle$$

$$\hat{H}_{\text{field}} |n\rangle = n\hbar\omega |n\rangle$$



with  $\hat{H}_I$  atom states now coupled to field states. Thence  $|i\rangle = |a\rangle |n\rangle$

↑ initial ground state



Since 1 photon carries  $1\hbar$  angular momentum parity conserved under  $\hat{d}_{if}$

$$H_{f,i}^I \equiv \langle f_i | \hat{H}^I | i \rangle = \langle b | \langle n-1 | [-\vec{E}_0 \cdot \hat{d} (\hat{a} - \hat{a}^\dagger)] | n \rangle | a \rangle$$

$$= -\vec{E}_0 \cdot \langle b | \hat{d} | a \rangle \langle n-1 | \hat{a} - \hat{a}^\dagger | n \rangle$$

$$= -\vec{E}_0 \cdot \vec{d}_{ba} \left[ \frac{\langle n-1 | \hat{a} | n \rangle}{\sqrt{n} \langle n-1 | n-1 \rangle} - \frac{\langle n-1 | \hat{a}^\dagger | n \rangle}{\sqrt{n+1} \langle n-1 | n+1 \rangle} \right]$$

$$H_{f,i}^{\pm} = -\vec{E}_0 \cdot \vec{d}_{ba} \sqrt{n} \quad \text{absorption}$$

Similarly

$$\begin{aligned} H_{f_2,i}^{\pm} &= \langle f_2 | \hat{H}^{\pm} | i \rangle \\ &= \langle b | \langle n+1 | [-\vec{E}_0 \cdot \hat{d} (\hat{a} - \hat{a}^{\dagger})] | n \rangle | a \rangle \\ &= -\vec{E}_0 \cdot \langle b | \hat{d} | a \rangle \langle n+1 | \hat{a} - \hat{a}^{\dagger} | n \rangle \\ &= -\vec{E}_0 \cdot \vec{d}_{ba} \left[ \sqrt{n} \langle n+1 | n \rangle - \sqrt{n+1} \langle n+1 | n+1 \rangle \right] \end{aligned}$$

$$H_{f_2,i}^{\pm} = \begin{array}{c} \sqrt{n+1} \\ \uparrow \\ \vec{E}_0 \cdot \vec{d}_{ba} \end{array} \quad \text{emission}$$

spontaneous  
stimulated

Loosely the probability of absorption and emission are

$$\begin{aligned} \frac{1}{E} P_{f,i}^{\pm} &= \frac{1}{E} |H_{f,i}^{\pm}|^2 = \frac{1}{E} |\vec{E}_0 \cdot \vec{d}_{ba}|^2 (n) = W_{f,i} \quad \text{absorption rate} \\ \frac{1}{E} P_{f_2,i}^{\pm} &= \frac{1}{E} |H_{f_2,i}^{\pm}|^2 = \frac{1}{E} |\vec{E}_0 \cdot \vec{d}_{ba}|^2 (n+1) = W_{f_2,i} \quad \text{emission rate} \end{aligned}$$

↑ ↑ spontaneous  
stimulated

Hence

$$\left. \begin{array}{l} \text{Stimulated Emission} \propto n \\ \text{Stimulated Absorption} \propto n \end{array} \right\} \text{Reciprocal / Classical} \\ \text{Detailed Balance}$$

$$\text{spontaneous Emission} \propto 1 \quad \left. \right\} \text{occurs when } n=0!$$

Stimulated processes proportional to Intensity.

Spontaneous emission occurs in Vacuum.

This is why an excited atom in empty space radiates!

Classical Analog: Radiation Reaction Theory.

Quantum: Vacuum Fluctuations "stimulate" Decay.

Einstein first pointed  $W_{fi} \neq W_{if}$  if Planck B.B. is correct. Added spontaneous emission in by hand 1905

$$\frac{W_{f_{2i}}^{\text{emission}}}{W_{f_{1i}}^{\text{absorption}}} = \frac{n+1}{n} \xrightarrow{n \rightarrow \infty} 1$$

For large  $n \rightarrow \infty$  we recover classical reciprocity.

Let  $\hbar\omega_a = E_a$  and  $\hbar\omega_b = E_b$  Atomic  $E_n = \hbar\omega_n = n\hbar\omega$   
 $\hbar\omega = E_0$  field  $\Rightarrow E_n^\pm = \hbar\omega_n^\pm = \hbar(n \pm 1)\omega$

With only  $|a\rangle, |b\rangle$  atom and  $|n\rangle$  field we have in general

$$|\Psi(t)\rangle = |\Psi_{\text{atom}}\rangle |\Psi_{\text{field}}\rangle$$

$$= C_i(t) |a\rangle |n\rangle e^{-i(\omega_a + \omega_n)t}$$

$$+ C_{f_1}(t) |b\rangle |n-1\rangle e^{-i(\omega_b + \omega_{n-1})t} \quad \text{ABS}$$

$$+ C_{f_2}(t) |b\rangle |n+1\rangle e^{-i(\omega_b + \omega_{n+1})t} \quad \text{EMIS}$$

General Solution:

If B.C.  $|\Psi(0)\rangle = |a\rangle |n\rangle \Rightarrow C_i(0) = 1 ; C_{f_{1,2}}(0) = 0$

we use Eq. 4.28

$$C_{f_1}(t) = -i\gamma \int_0^t dt' H_{f_{1i}}^I e^{i\omega_{f_{1i}} t'}$$

$$C_{f_2}(t) = -i\gamma \int_0^t dt' H_{f_{2i}}^I e^{i\omega_{f_{2i}} t'}$$

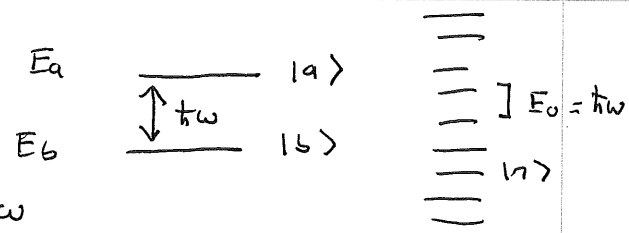
$$\omega_{f_{1i}} = \omega_{f_{1i}} = (\omega_b + \omega_{n-1}) - (\omega_a + \omega_n) = \omega_{ba} + \omega = -\Omega_{ba}^+$$

$$\omega_{f_{2i}} = (\omega_b + \omega_{n+1}) - (\omega_a + \omega_n) = \omega_{ba} - \omega = -\Omega_{ba}^-$$



# Laser Rate Equations (Einstein)

Suppose a collection of two level atoms are resonant



$\Delta = 0$  with  $\omega_{field} = \omega_{ab} = \omega$

Let  $N_a \equiv \#$  atoms in  $|a\rangle$  (population)

$N_b \equiv \#$  atoms in  $|b\rangle$  "

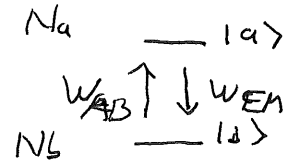
Let  $W_{em}^{EM} =$  emission rate  $= P_{em}/t \propto (n+1)$

$W_{abs}^{AB} =$  absorption rate  $= P_{abs}/t \propto (n)$

The rate equations are  $N_a = N_a(t)$ ;  $N_b = N_b(t)$

$$\frac{dN_a}{dt} = -W_{em}^{EM} N_a + N_b W_{abs}^{AB} \quad (II)$$

$$\frac{dN_b}{dt} = -W_{abs}^{AB} N_b + W_{em}^{EM} N_a \quad (I)$$



Note  $\frac{W_{em}^{EM}}{W_{abs}^{AB}} = \frac{n+1}{n}$  so there is an asymmetry.

Einstein now assumes thermal equilibrium

$$\Rightarrow \dot{N}_a = \dot{N}_b \equiv 0$$

$$\Rightarrow N_a W_{em}^{EM} = N_b W_{abs}^{AB} \Rightarrow \boxed{\frac{N_b}{N_a} = \frac{W_{em}^{EM}}{W_{abs}^{AB}} = \frac{n+1}{n}} \quad \downarrow \text{S.E.}$$

But from thermo  $\frac{N_b}{N_a} = e^{(E_a - E_b)/kT} \equiv e^{hw/kT}$  //

solve

$$\frac{n+1}{n} = e^{hw/kT} \Rightarrow \boxed{n = \bar{n} = \frac{1}{e^{hw/kT} - 1}} \quad \uparrow \text{S.E.} \quad \text{!} \text{!} \text{!} \text{!} \text{!}$$

Hence spontaneous emission / vacuum / rad. react  
is required to get Planck!

for  $n \gg 1$   $\frac{n+1}{n} \approx 1 \Rightarrow$  Boltzman.

Einstein did not have quantum optics

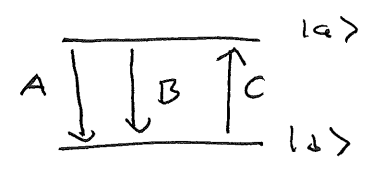
Einstein used spectral energy ~~U~~  $U(\omega) \propto I \propto n$

Introduced  $A, B, C$

and explicitly states  $A$

is spontaneous rate independent

of  $U(\omega)$ .  $B$  is stimulated emission,  $C$  absorption



$$\dot{N}_a = - \overset{\substack{SE \\ \downarrow}}{A} N_a - \overset{\substack{SE \\ \downarrow}}{B U} N_a + \overset{\substack{SA \\ \downarrow}}{C U} N_b$$

$$\dot{N}_b = - \overset{\substack{SA \\ \uparrow}}{C U} N_a + \overset{\substack{SE \\ \downarrow}}{A + B U} N_a$$

again  $\dot{N}_a = \dot{N}_b = 0$

$$\Rightarrow U(\omega) = \frac{A}{C e^x - B} \quad x = \frac{h\omega}{kT}$$

To recover Planck  $C = B$  and  $\frac{A}{B} = \frac{h\omega^3}{\pi^2 c^3}$

22-141 50 SHEETS  
 22-142 100 SHEETS  
 22-143 200 SHEETS  
 AMPAZ

Let us assume  $N_a^{(0)} = N$  ;  $N_b^{(0)} = 0$

~~⊗~~  $|a\rangle$

called inversion. All atoms up!

—  $|b\rangle$

$$\dot{N}_a = -W_E N_a + N_b W_A ; \quad \dot{N}_b = -W_A N_b + W_E N_a$$

$$\Rightarrow \ddot{N}_a = -W_E \dot{N}_a + [-W_A N_b + W_E N_a]$$

For short times  $\left\{ \begin{array}{l} N_b \approx 0 \\ \cancel{N_a = N} \end{array} \right\}$  since no population in  $|b\rangle$   
 $N_a = N$

$$\Rightarrow \ddot{N}_a \approx -W_E (\dot{N}_a \neq N_a)$$

$$\Rightarrow \ddot{N}_a + W_E \dot{N}_a \neq W_E N_a \approx 0$$

↗ No role in dynamics

Also

$$\Rightarrow \dot{N}_a = -W_E N_a$$

$$\Rightarrow \boxed{N_a(t) = e^{-g(n+1)t} N_a(0)}$$

$$g = 4\gamma^2 |\vec{d}_{ba} \cdot \vec{E}_0|^2 = \text{gain laser gain}$$

Atoms decay from  $|a\rangle \rightarrow |b\rangle$  at a rate proportional to  $(n+1)$ .

The more photons in the field  
 the more quickly atoms decay  $\Rightarrow$  photons  
 accumulate Exp. Fast. Light Amp. Stim. Emiss.