

Geometric series for Ch 2.5

Theorem ΘI : If $|r| < 1$ Then $\sum_{n=0}^{\infty} r^n = \frac{1}{1-r}$

Proof: If $|r| < 1$ series is absolutely convergent and we may manipulate it term by term

$$\text{Let } S_0 \equiv \sum_{n=0}^{\infty} r^n = 1 + r + r^2 + \dots$$

$$\Rightarrow r S_0 = \sum_{n=0}^{\infty} r^{n+1} = r + r^2 + \dots$$

$$\Rightarrow S_0 - r S_0 = 1 \quad \Rightarrow \quad S_0 = \frac{1}{1-r} \quad \therefore$$

Corollary ΓI : If $0 \leq r < 1$ $P_n \equiv (1-r)r^n$

is a normalized probability distribution.

Proof $0 \leq r < 1 \Rightarrow 0 \leq r^n < 1$ and $0 \leq 1-r < 1 \Rightarrow 0 \leq P_n < 1$

$$\text{Further } \sum_{n=0}^{\infty} P_n = (1-r) \sum_{n=0}^{\infty} r^n = \frac{(1-r)}{(1-r)} = 1 \quad \therefore$$

ΘII : $|r| < 1 \Rightarrow \sum_{n=0}^{\infty} n r^n = \frac{r}{(1-r)^2} \equiv S_1$

$$\text{Proof } \frac{d S_0}{dr} = \frac{d}{dr} \sum_{n=0}^{\infty} r^n = \sum_{n=0}^{\infty} n r^{n-1} \equiv \frac{d}{dr} \left(\frac{1}{1-r} \right) = \frac{1}{(1-r)^2}$$

$$\Rightarrow r \frac{d S_0}{dr} = \sum_{n=0}^{\infty} n r^n \equiv S_1 = \frac{r}{(1-r)^2} \quad \therefore$$

ΓII : If $0 \leq r < 1$ Then $\langle n \rangle = \bar{n} \equiv \sum_{n=0}^{\infty} n P_n = \frac{r}{(1-r)}$

ΘIII : $|r| < 1 \Rightarrow \sum_{n=0}^{\infty} n^2 r^n \equiv S_2 = \frac{r(1+r)}{(1-r)^3}$

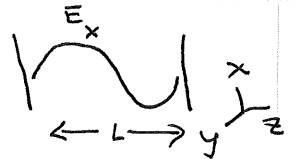
$$\text{Proof } \frac{d}{dr} \sum_{n=0}^{\infty} n r^n = \sum_{n=0}^{\infty} n^2 r^{n-1} \equiv \frac{d}{dr} \left[\frac{r}{(1-r)^2} \right] = \dots = \frac{1+r}{(1-r)^3} \quad \therefore$$

$$\Rightarrow S_2 = \frac{r(1+r)}{(1-r)^3}$$

ΓIII : $0 \leq r < 1 \Rightarrow \langle n^2 \rangle \equiv \overline{n^2} \equiv \sum_{n=0}^{\infty} n^2 P_n = \frac{r(1+r)}{(1-r)^2}$

Proof: Use ΘIII , multiply by $(1-r)$ and note $P_n = (1-r)r^n$

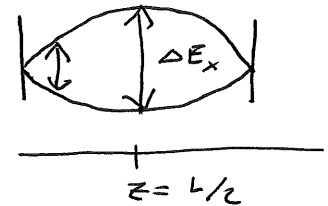
For single mode in a cavity recall electric field that $\Delta E_x^{(n)} = \sqrt{2} \epsilon_0 (\sin kz) \sqrt{n+1/2}$



For n photons and $E_0 = (\hbar\omega/\epsilon_0 V)^{1/2}$ is the single-photon Electric Field amplitude. Hence the electric field fluctuates vacuum $n=0$ $\Delta E_x^{(0)} = E_0 \sin kz$

For a cavity $L = \lambda/2 \Rightarrow k = \frac{2\pi}{\lambda} = \frac{\pi}{L}$ and

$\sin(\frac{\pi}{L}z)$ is maximum at $z = L/2$



The zero point energy per mode \vec{k}, s

is $E_{\vec{k}s} = \frac{\hbar}{2} \omega_k$

In empty space summing over all modes:

$$E_{ZPE} = \sum_{\vec{k}s} E_{\vec{k}s} = \frac{\hbar}{2} \sum_{\vec{k}s} \omega_k \rightarrow \frac{\hbar}{2} \int \omega_k \rho(\omega) d\omega$$

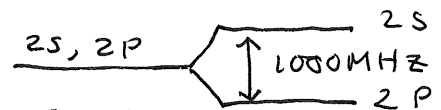
$$= \frac{\hbar}{2} \frac{V}{\pi^2 c^3} \int_0^\infty \omega^3 d\omega = \frac{\hbar V}{8\pi^2 c^3} \lim_{\omega \rightarrow \infty} \omega^4 = \infty \text{ Diverges Quite Fast!}$$

Nevertheless ZPE can be viewed as causing real effects such as Lamb shift and Casimir effect,

Lamb shift

In Dirac Eq. for hydrogen the 2S and 2P

excited states are degenerate.



But interaction with quantized $E \& B$ Field

lifts the degeneracy, splitting the levels. Observed by Lamb in 1947, estimated in theory by Bethe 1948.

Also in 1948 Welton provided ZPE interpretation.

$$E_{\text{ZPE}}(\omega) = \frac{1}{2} \hbar \omega \quad \text{per mode}$$

but in free space # modes from $\omega \rightarrow \omega + d\omega$ is DOM

$$\text{DOM} = \int_V \rho(\omega) d\omega = \frac{V \omega^2}{\pi^2 c^3} d\omega \quad \text{hence over this range}$$

$$E_{\text{ZPE}} = \text{DOM} \times \frac{1}{2} \hbar \omega = \frac{V \hbar}{2 \pi^2 c^3} \omega^3 d\omega$$

Now atom only couples to electric field energy

$$E_{\text{ZPE}} = \frac{1}{2} \int_V dV (\epsilon_0 E^2 + \frac{1}{\mu_0} B^2) \quad \xrightarrow{\text{neglect}} = \frac{\epsilon_0 V}{2} E^2(\omega)$$

where E is electric field at frequency ω .

$$\Rightarrow \frac{V \hbar}{2 \pi^2 c^3} \omega^3 d\omega = \frac{\epsilon_0 V}{2} E^2(\omega)$$

$$\Rightarrow \boxed{E^2(\omega) = \frac{\hbar \omega^3 d\omega}{\pi^2 c^3 \epsilon_0}} \quad \text{were going to do this in SI units unlike bohr}$$

so we assume atom is bathed in a field of this strength.

We also assume the coulomb potential $V(r) = -\frac{1}{4\pi\epsilon_0} \frac{e^2}{r}$

is modified by the vacuum field. To estimate how much:

$$\Delta V(x, y, z) = \Delta x \frac{\partial V}{\partial x} + \Delta y \frac{\partial V}{\partial y} + \Delta z \frac{\partial V}{\partial z} + \frac{1}{2} \Delta x^2 \frac{\partial^2 V}{\partial x^2} + \frac{1}{2} \Delta y^2 \frac{\partial^2 V}{\partial y^2} + \frac{1}{2} \Delta z^2 \frac{\partial^2 V}{\partial z^2} + \dots$$

which is a 3D Taylor's series expansion.

OK recall $\langle E \rangle = 0$ for the vacuum electric field

and only $\Delta E^2 \neq 0$ (see Eq. 2.44 and Eq. 2.45).

since for an electron $\Delta x, \Delta y, \Delta z \propto \langle E \rangle = 0$ only

second term survives. If isotropic free space

$$\Delta x^2 = \Delta y^2 = \Delta z^2 = \Delta r^2 / 3 \quad \text{and}$$

$$\Delta V \approx \frac{1}{2} \cdot \frac{1}{3} \Delta r^2 \left[\frac{\partial^2 V}{\partial x^2} + \frac{\partial^2 V}{\partial y^2} + \frac{\partial^2 V}{\partial z^2} \right] = \boxed{\frac{1}{6} \Delta r^2 \nabla^2 V}$$

If $V(r) = \frac{-1}{4\pi\epsilon_0} \frac{e^2}{r}$ use NRL Eq. 25 & 24

$$\vec{\nabla} \cdot \vec{\nabla} (1/r) = \vec{\nabla} \cdot \left[-\frac{\vec{r}}{r^3} \right] = -4\pi \delta(\vec{r})$$

$$\Rightarrow \Delta V = \frac{1}{6} \Delta r^2 \left[\frac{4\pi e^2}{4\pi\epsilon_0} \right] \delta(\vec{r}) = \frac{1}{6} \frac{e^2}{\epsilon_0} \Delta r^2 \delta(\vec{r})$$

For a hydrogen state $|nlm\rangle = \Psi_{nlm}(\vec{r})$

the wavefunction vanishes at $\vec{r}=0$ unless $l=m=0$ (s-state).

Hence the energy shift will be

$$\begin{aligned} \Delta E &= \langle nlm | \Delta V | nlm \rangle \\ &= \frac{1}{6} \frac{e^2}{\epsilon_0} \Delta r^2 \int dV \Psi_{nlm}^*(\vec{r}) \delta(\vec{r}) \Psi_{nlm}(\vec{r}) \\ &= \frac{1}{6} \frac{e^2}{\epsilon_0} \Delta r^2 |\Psi_{n00}(0)|^2 \end{aligned}$$

One can prove $|\Psi_{n00}(0)|^2 = \frac{1}{\pi n^3 a_0^3}$

where $a_0 = \frac{\epsilon_0 \hbar^2}{\pi m e^2} = \frac{\pi \epsilon_0 \hbar^2}{m e^2}$ is Bohr radius
NRL pg 14

$$\Rightarrow \Delta E_{\text{SHIFT}}^{(n)} = \frac{e^2}{6\pi\epsilon_0 a_0^3 n^3} \Delta r^2 \quad \text{for the state } |n00\rangle$$

Hence only s-STATES ($l=0$) experience a shift.

To compute Δr^2 we assume the electron experiences the fluctuating electric field $\mathcal{E}^2(\omega)$

$$F = ma \Rightarrow e \mathcal{E}(\omega) \stackrel{\downarrow, t}{\approx} m \frac{d^2}{dt^2} \Delta r(\omega)$$

$$\text{Let } \mathcal{E}(\omega, t) = \mathcal{E}(\omega) e^{i\omega t}$$

$$\Rightarrow \frac{d^2 \Delta r(\omega, t)}{dt^2} = \frac{e \mathcal{E}(\omega)}{m} e^{i\omega t}$$

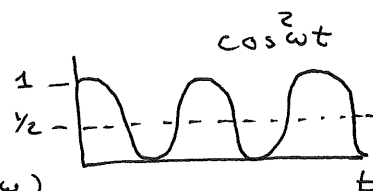
$$\Rightarrow \Delta r(\omega, t) = -\frac{e \mathcal{E}(\omega)}{m \omega^2} e^{i\omega t}$$

Note only the real part of the electric field can contribute to the electron motion: $e^{i\omega t} = \cos\omega t + i\sin\omega t$

$$\Rightarrow \Delta r^2(\omega, t) = \frac{e^2 \mathcal{E}^2(\omega)}{m^2 \omega^4} \cos^2 \omega t$$

The time average of $\cos^2 \omega t = 1/2$

$$\Rightarrow \langle \Delta r^2(\omega, t) \rangle_{\text{Time}} = \frac{e^2 \mathcal{E}^2(\omega)}{2 \omega^4} = \Delta r^2(\omega)$$



We can now use expression $\mathcal{E}^2(\omega) = \frac{\hbar \omega^3}{\pi^2 c^3 \epsilon_0} d\omega$

$$\Rightarrow \boxed{\Delta r^2(\omega) = \frac{1}{2} \frac{\hbar e^2}{\pi^2 c^3 \epsilon_0 m^2} \frac{d\omega}{\omega}}$$

$$\Rightarrow \Delta E_n^{\text{SHIFT}}(\omega) = \left(\frac{e^2}{6\pi \epsilon_0 n^3 a_0^3} \right) \left(\frac{\hbar e^2}{2\pi^2 c^3 \epsilon_0 m^2} \right) \frac{d\omega}{\omega}$$

Recall pg 14 NRL that Fine structure constant $\alpha = \frac{e^2}{4\pi \epsilon_0 \hbar c}$

and Compton wavelength $\lambda_c = \frac{\hbar}{mc}$

$$\begin{aligned} \Rightarrow \Delta E_n(\omega) &= \frac{\hbar e^4}{12\pi^3 \epsilon_0^2 c^3 m^2 n^3 a_0^3} \frac{d\omega}{\omega} \\ &= \frac{\hbar}{12\pi^3 \epsilon_0^2 c^3 n^3 a_0^3} \left[\frac{\hbar^2}{m^2 c^2} \frac{c^2}{\hbar^2} \right] \left[\frac{e^4 (4\pi \epsilon_0 \hbar c)^2}{(4\pi \epsilon_0 \hbar c)^2} \right] \frac{d\omega}{\omega} \\ &= \frac{\hbar}{12\pi^3 \epsilon_0^2 c^3 n^3 a_0^3} \frac{c^2}{\hbar^2} \alpha^2 \lambda_c^2 (4\pi \epsilon_0 \hbar c)^2 \frac{d\omega}{\omega} \\ &= \frac{16\pi^2 \epsilon_0^2 \hbar^3 c^4}{12\pi^3 \epsilon_0^2 \hbar^2 c^3 n^3 a_0^3} \alpha^2 \lambda_c^2 \frac{d\omega}{\omega} \\ &= \frac{4}{3\pi} \frac{\hbar c}{n^3 a_0^3} \alpha^2 \lambda_c^2 \frac{d\omega}{\omega} \quad \left(\hbar = \frac{\hbar}{2\pi} \right) \\ &= \frac{2}{3} \alpha^2 \lambda_c^2 \frac{\hbar c}{n^3 a_0^3} \frac{d\nu}{\nu} \quad \left(\frac{d\omega}{\omega} = \frac{d\nu}{\nu} \right) \\ &= \text{Eq. 2.168} \quad \text{(:)} \quad \left(2\pi\nu = \omega \right) \end{aligned}$$

The total shift is obtained by integrating over all frequencies but $\int_0^{\infty} \frac{d\omega}{\omega} = \ln(\infty) - \ln(0) = \infty + \infty = \infty$.

So we must cut off the integral.

$$\int_0^{\infty} \frac{d\omega}{\omega} \rightarrow \int_{\Omega_L}^{\Omega_U} \frac{d\omega}{\omega}$$

At $\Omega_L \rightarrow 0$ the rapidly orbiting electron is out of resonance

$$\begin{aligned} \text{so } \Omega_L &= \omega_0 = 2\pi\nu_0 = \left(\frac{e}{4\pi\epsilon_0}\right)^2 \frac{2\pi}{\hbar a_0^3 n^3} && \text{orbit frequency in } n^{\text{th}} \text{ state} \\ &= \left(\frac{e^2}{4\pi\epsilon_0 \hbar c}\right) \frac{\hbar c}{4\pi\epsilon_0} \frac{2\pi}{\hbar a_0^3 n^3} = \alpha \left(\frac{c}{2\epsilon_0 a_0^3 n^3}\right) \end{aligned}$$

For upper limit $\hbar\omega \leq mc^2$ to avoid relativity

$$\Rightarrow \boxed{\Omega_U = mc^2 / \hbar}$$

Hence $\Delta E_n \equiv \int_{\Omega_L}^{\Omega_U} \Delta E_n(\omega) d\omega$

$$= \frac{4}{3\pi} \frac{\hbar c \alpha^2 \hbar^2}{n^3 a_0^3} \int_{\Omega_L}^{\Omega_U} \frac{d\omega}{\omega} = \frac{4}{3\pi} \frac{\hbar c \alpha^2 \hbar^2}{n^3 a_0^3} \ln \left[\frac{\Omega_U}{\Omega_L} \right]$$

$$= \frac{4}{3\pi} \frac{\hbar c}{n^3} \frac{\alpha^2 \hbar^2}{n^3 a_0^3} \ln \left[\frac{mc^2}{\hbar\omega_0} \right]$$

For 2s $n=2 \Rightarrow \hbar\omega_0 = \frac{\hbar c \alpha}{16\epsilon_0 a_0^3}$

$$\Delta E_{2s} = \frac{1}{6\pi} \frac{\hbar c \alpha^2 \hbar^2}{a_0^3} \ln \left[\frac{mc^2}{\hbar\omega_0} \right]$$

$$= 9.06177 \times 10^{-25} \text{ J}$$

$$\frac{\Delta E_{2s}}{\hbar} = \boxed{1368 \text{ MHz}}$$

$$\frac{\Delta E_{2s}}{\hbar} (\text{exp}) = 1058 \text{ MHz}$$

error of 30%